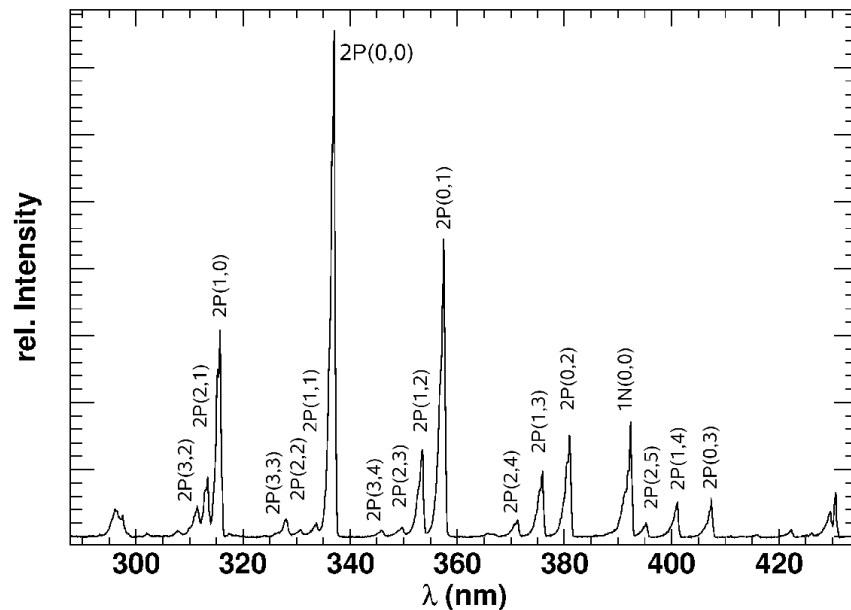


Measurement of the Air Fluorescence Yield by *AIRFLY*



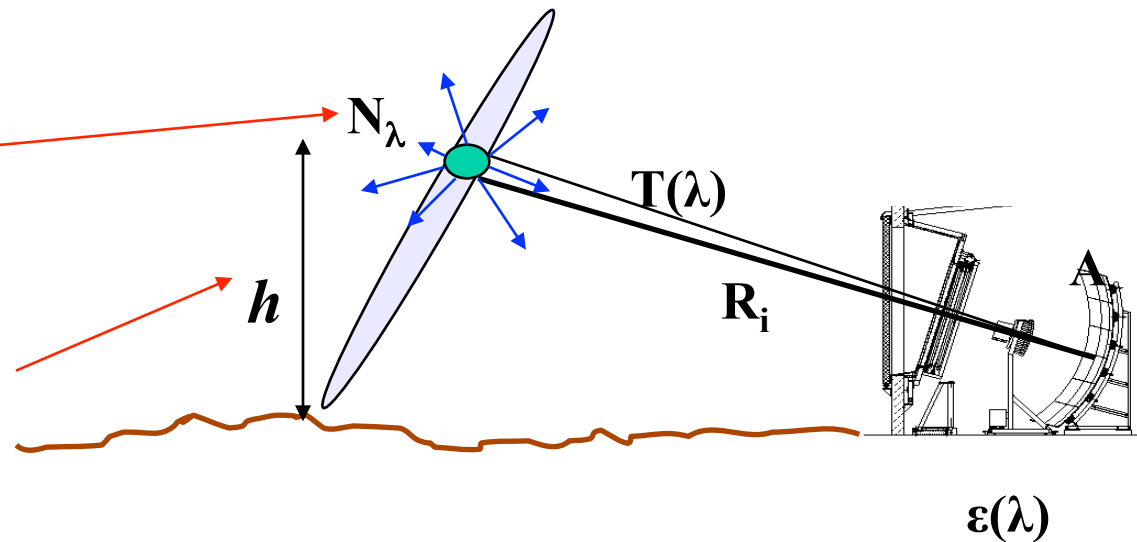
Paolo Privitera



The Enrico Fermi Institute
The Kavli Institute for Cosmological Physics
for the AIRFLY Collab.

UHECR Fluorescence Detection and Fluorescence Yield

- Fluorescence Spectrum
- p , T and humidity dependence
- Proportionality to energy deposit
- Absolute yield



$$N_\lambda^{\text{shower}}(h) = E_{\text{dep}}^{\text{shower}}(h) Y_{\text{air}}(\lambda, p, T)$$

$\Delta E/E = 22\%$ in Auger - largest syst. is 14% from absolute yield

AIRFLY strategy

- Precise measurement of fluorescence band intensities dependence on p, T, and H ([relative to 337 nm band](#))
- Measurement of the **absolute yield of 337 nm band**

$$Y_{\text{air}}(\lambda, p, T) = Y_{\text{air}}(337, p_0, T_0) I_{\lambda}(p_0, T_0) \frac{1 + \frac{p_0}{p'_{\text{air}}(\lambda, T_0)}}{1 + \frac{p}{p'_{\text{air}}(\lambda, T_0) \sqrt{\frac{T H_{\lambda}(T_0)}{T_0 H_{\lambda}(T)}}}} .$$

**Absolute yield
of 337 nm line
(photons/MeV)**

**Band intensity
relative to the
337 nm line**

**Quenching
reference
pressure**

**Temperature
dependence**

AIRFLY Collaboration: M. Ave^a, M. Bohacova^b,
 B. Buonomo^c, N. Busca^a, L. Cazon^a, S.D. Chemerisov^d,
 M.E. Conde^d, R.A. Crowell^d, P. Di Carlo^e, C. Di Giulio^f,
 M. Doubrava^g, A. Esposito^c, P. Facal^h, F.J. Franchini^d,
 J. Hörandelⁱ, M. Hrabovsky^b, M. Iarlori^e, T.E. Kasprzyk^d,
 B. Keilhauerⁱ, H. Klages^j, M. Kleifges^k, S. Kuhlmann^d,
 G. Mazzitelli^c, L. Nozka^b, A. Obermeierⁱ, M. Palatka^b,
 S. Petrera^e, P. Privitera^{f,*}, J. Ridky^b, V. Rizi^e,
 G. Rodriguez^f, F. Salamida^e, P. Schovanek^b, H. Spinka^d,
 E. Strazzeri^f, A. Ulrich^ℓ, Z.M. Yusof^d, V. Vacek^g,
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AIRFLY

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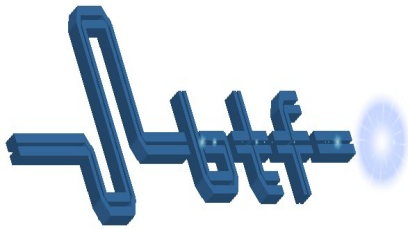
^k *Forschungszentrum Karlsruhe, Institut für Prozessdatenverarbeitung und Elektronik, Postfach 3640, D - 76021 Karlsruhe, Germany*

^ℓ *Physik Department E12, Technische Universität Muenchen, James Franck Str. 1, D-85748 Garching, Germany*

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+ Fermilab (C. Hojvat, F. Kuehn)

AIRFLY beams



Beam Test Facility INFN FRASCATI (50-750 MeV)

Chemistry Division Van de Graaff (0.5-3 MeV)



Advanced
Photon
Source

(6-30 KeV)



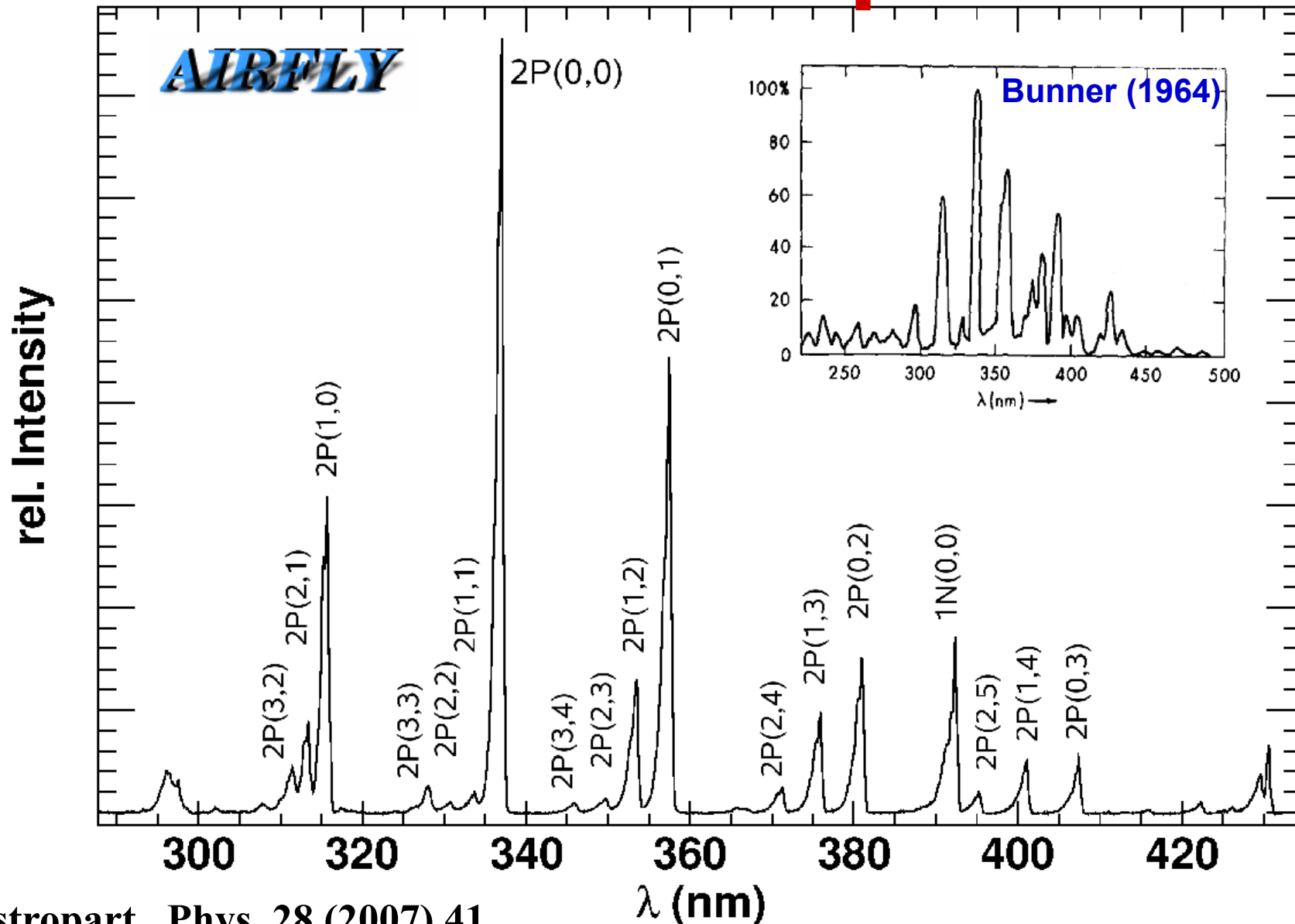
HEP Division
Advanced
Wakefield
Accelerator

(3 MeV-15 MeV)



Meson Test Beam (Up to 120 GeV)

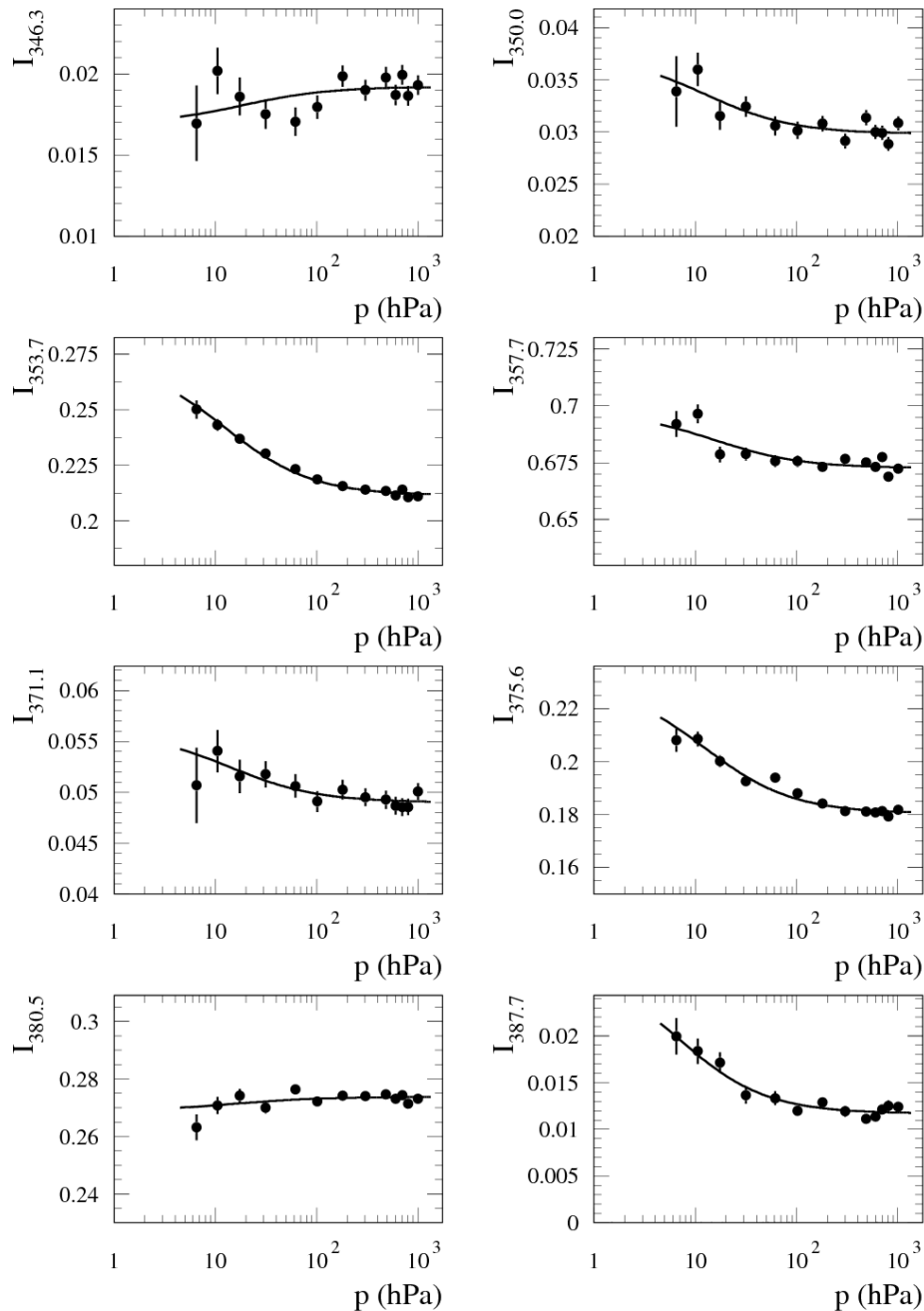
Air fluorescence spectrum



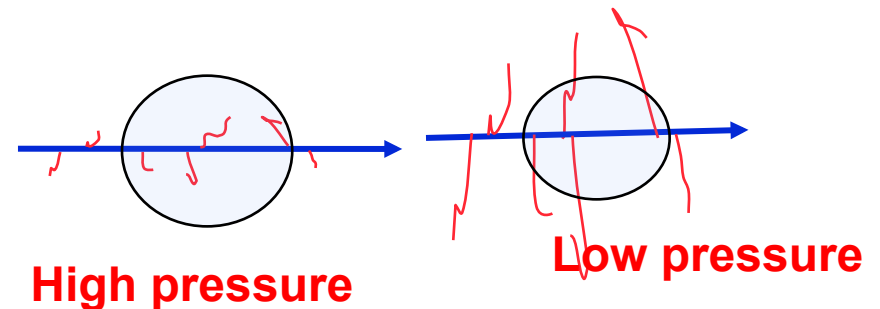
Astropart. Phys. 28 (2007) 41

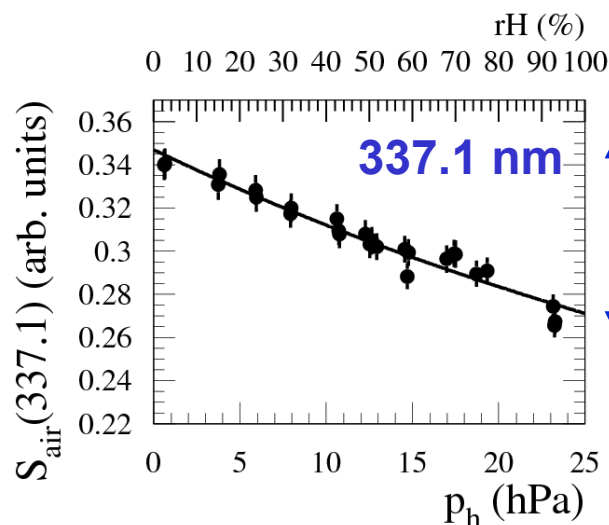
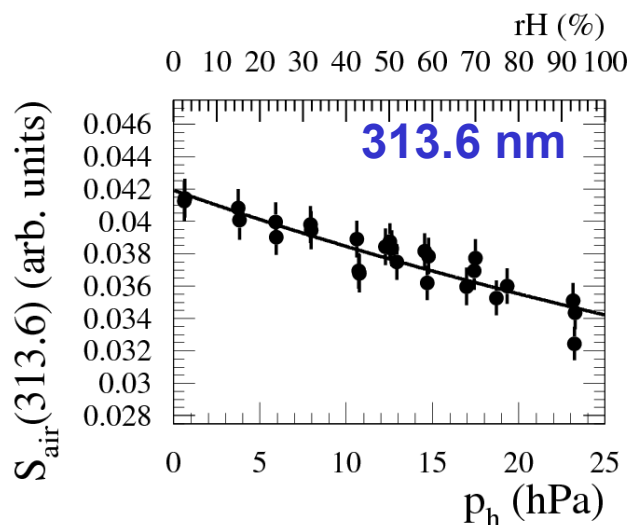
34 band intensities measured (relative to the 337 nm line)

Pressure dependence



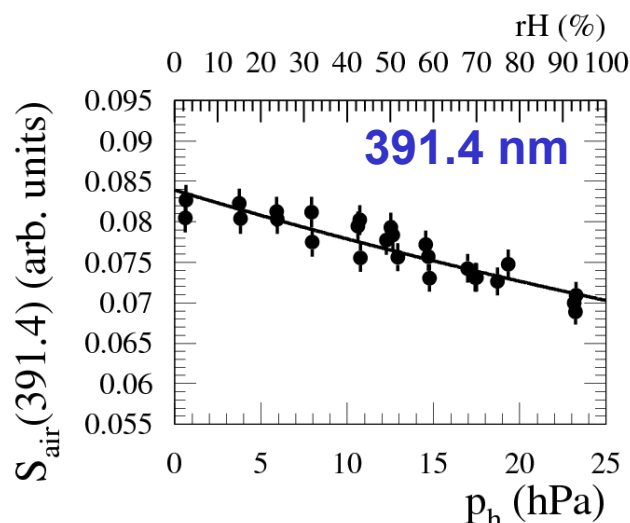
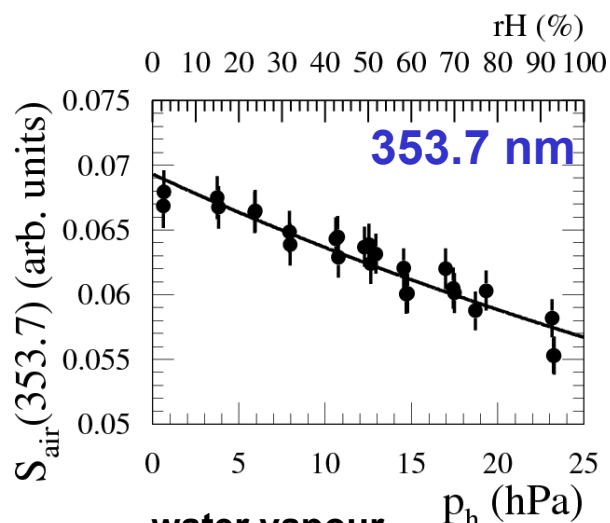
- AIRFLY measurement solved previous discrepancies due to neglecting secondary electrons escaping the field of view
(~all the fluorescence emission is produced by secondary electrons)





20 % !

**Humidity
dependence**



**water vapour
partial pressure**

$$\frac{1}{p'_{air}} \rightarrow \frac{1}{p'_{air}} \left(1 - \frac{p_h}{p} \right) + \frac{1}{p'_{H_2O}} \frac{p_h}{p}$$

2P(2,1)

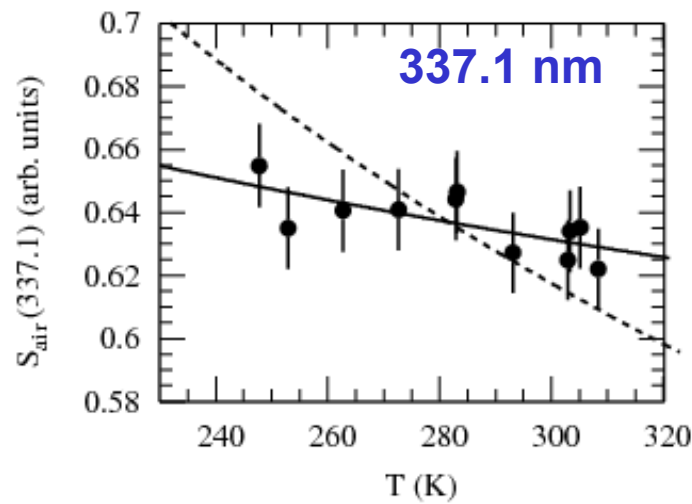
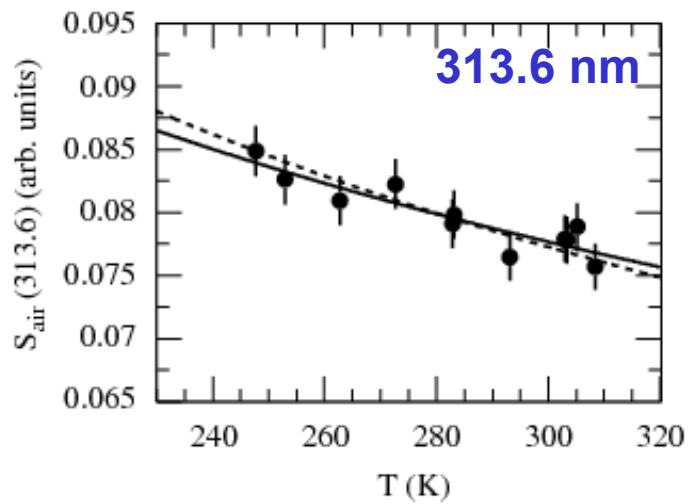
2P(0,0)

2P(1,2)

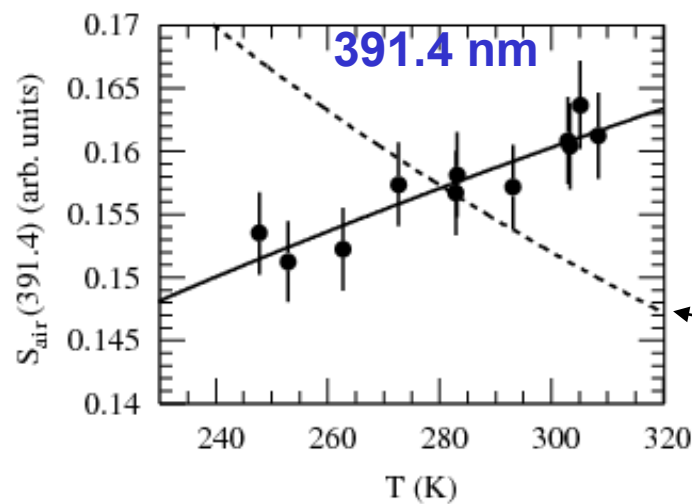
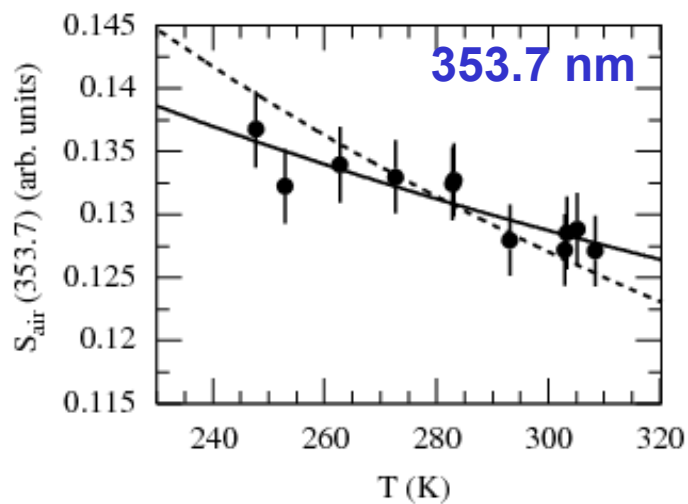
1N(0,0)

NIM A597 (2008) 50

λ (nm)	p'_{H_2O} (hPa)
313.6	1.21 ± 0.13
337.1	1.28 ± 0.08
353.7	1.27 ± 0.12
391.4	0.33 ± 0.03



RANGE [-40,30] °C



up to 10 % effect!

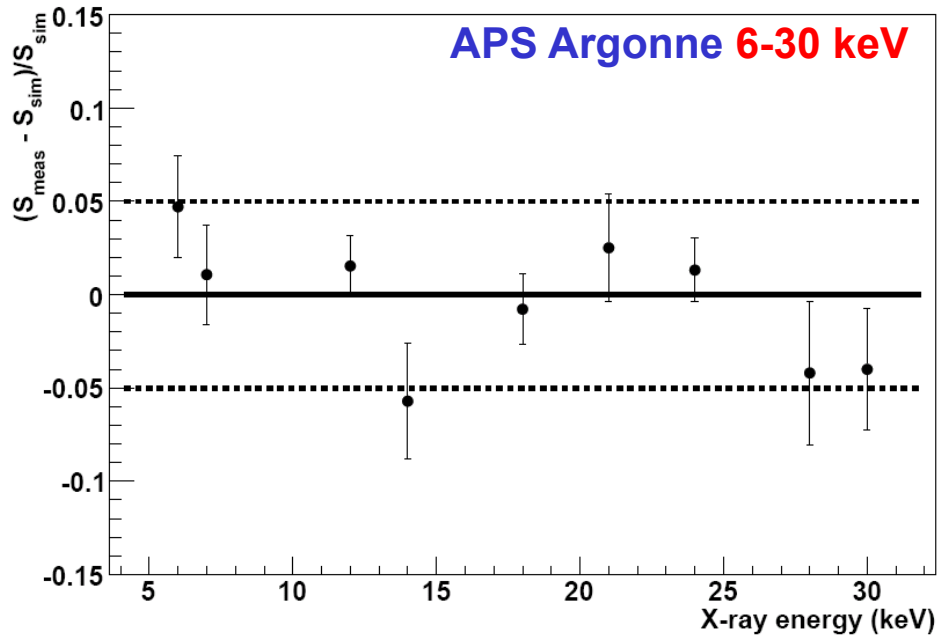
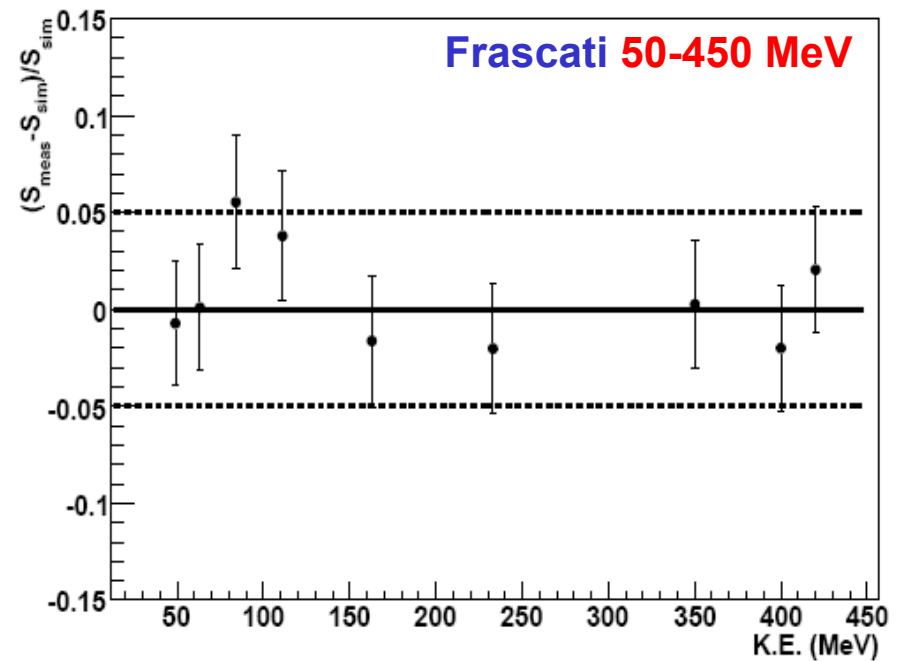
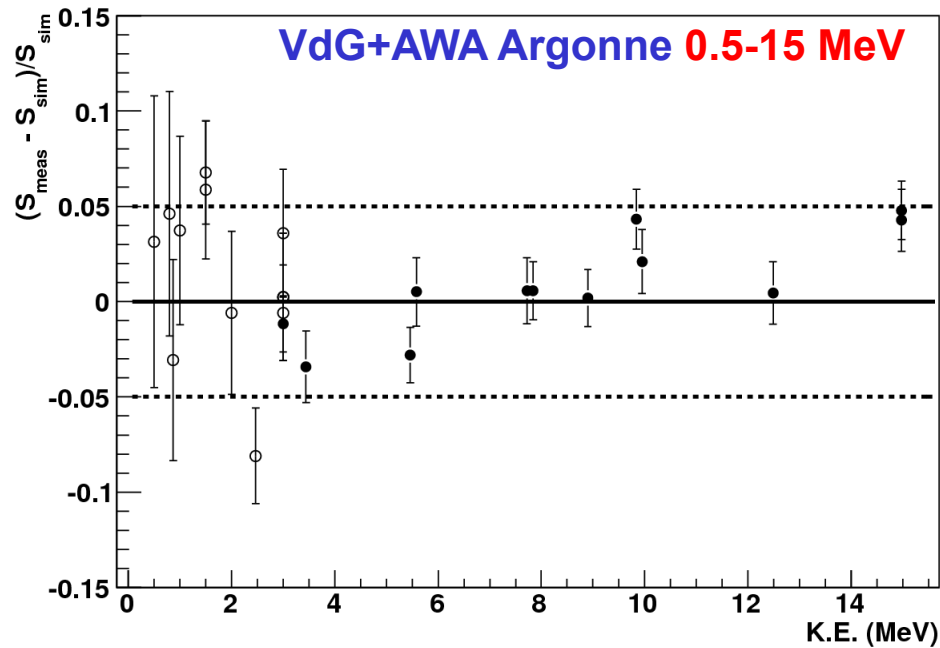
$\alpha = 0$
(assumed)

**Temperature
dependence**

$\alpha \neq 0$!

2P(2,1)
2P(0,0)
2P(1,2)
1N(0,0)

λ (nm)	α_λ
313.6	-0.09 ± 0.10
337.1	-0.36 ± 0.08
353.7	-0.21 ± 0.09
391.4	-0.80 ± 0.09



Energy dependence

NIM A597 (2008) 46

Proportionality of fluorescence
yield to E_{dep} at few %

Absolute yield of the 337 nm line

Main systematic comes from absolute QE of photon detector

AIRFLY: normalize to a well known process (Cherenkov emission) to cancel detector systematics

$$N_{337}(\text{fluor.}) = \text{FLY} \times \text{Geom}_{\text{fluor}} \times T_{\text{filter}} \times \text{QE}_{337}$$

measured (pointing to $N_{337}(\text{fluor.})$)

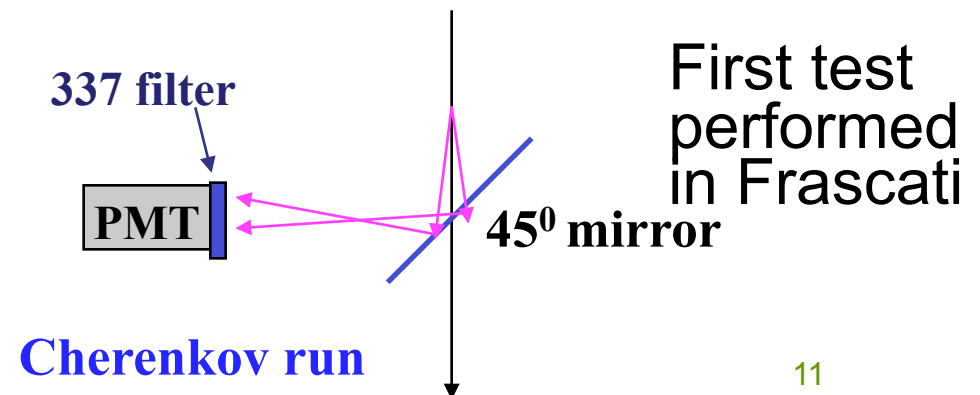
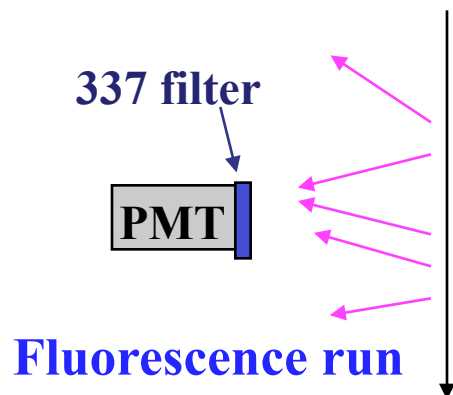
known (pointing to FLY)

MC (pointing to $\text{Geom}_{\text{fluor}}$)

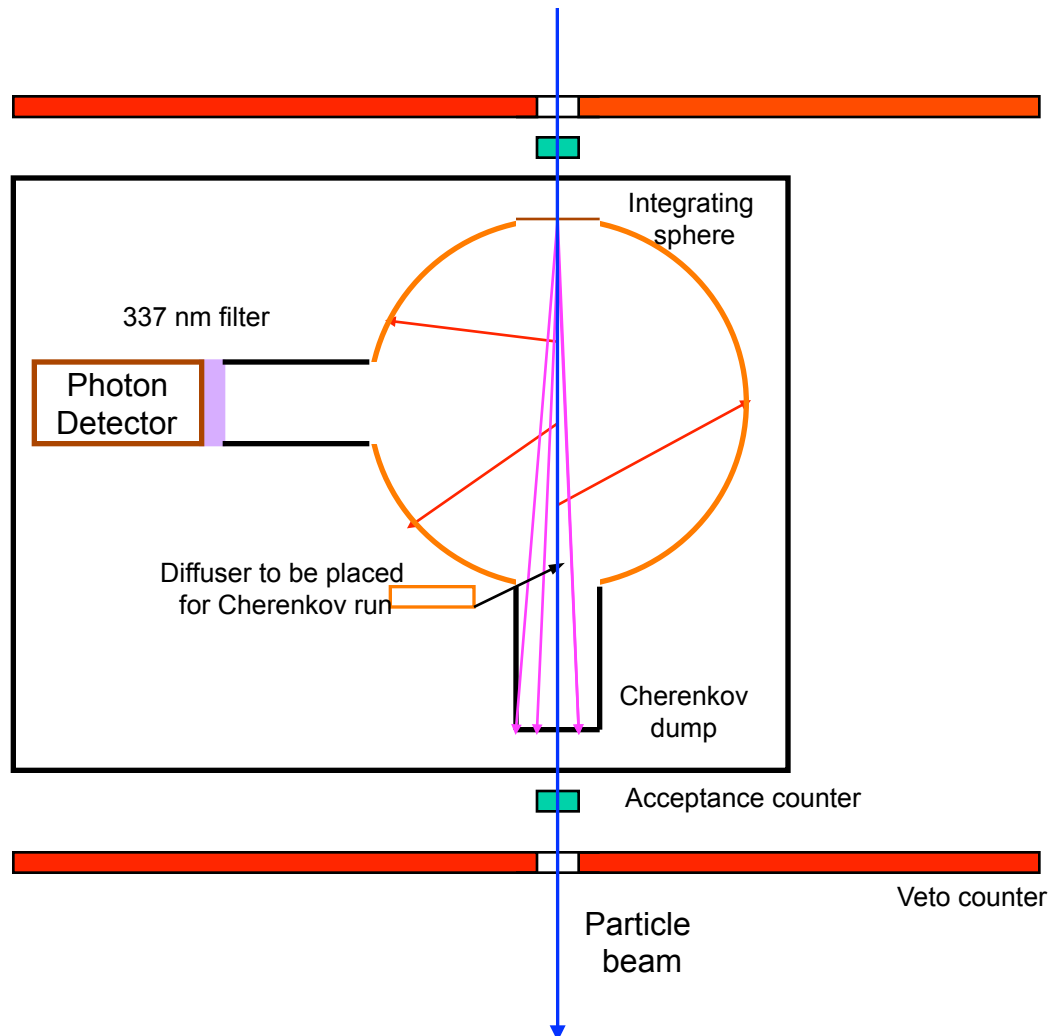
~cancel! (pointing to T_{filter} and QE_{337})

measured (pointing to QE_{337})

$$N_{337}(\text{Cher.}) = \text{CHY} \times \text{Geom}_{\text{cher}} \times T_{\text{filter}} \times \text{QE}_{337} \times \epsilon_{\text{mirr}}$$

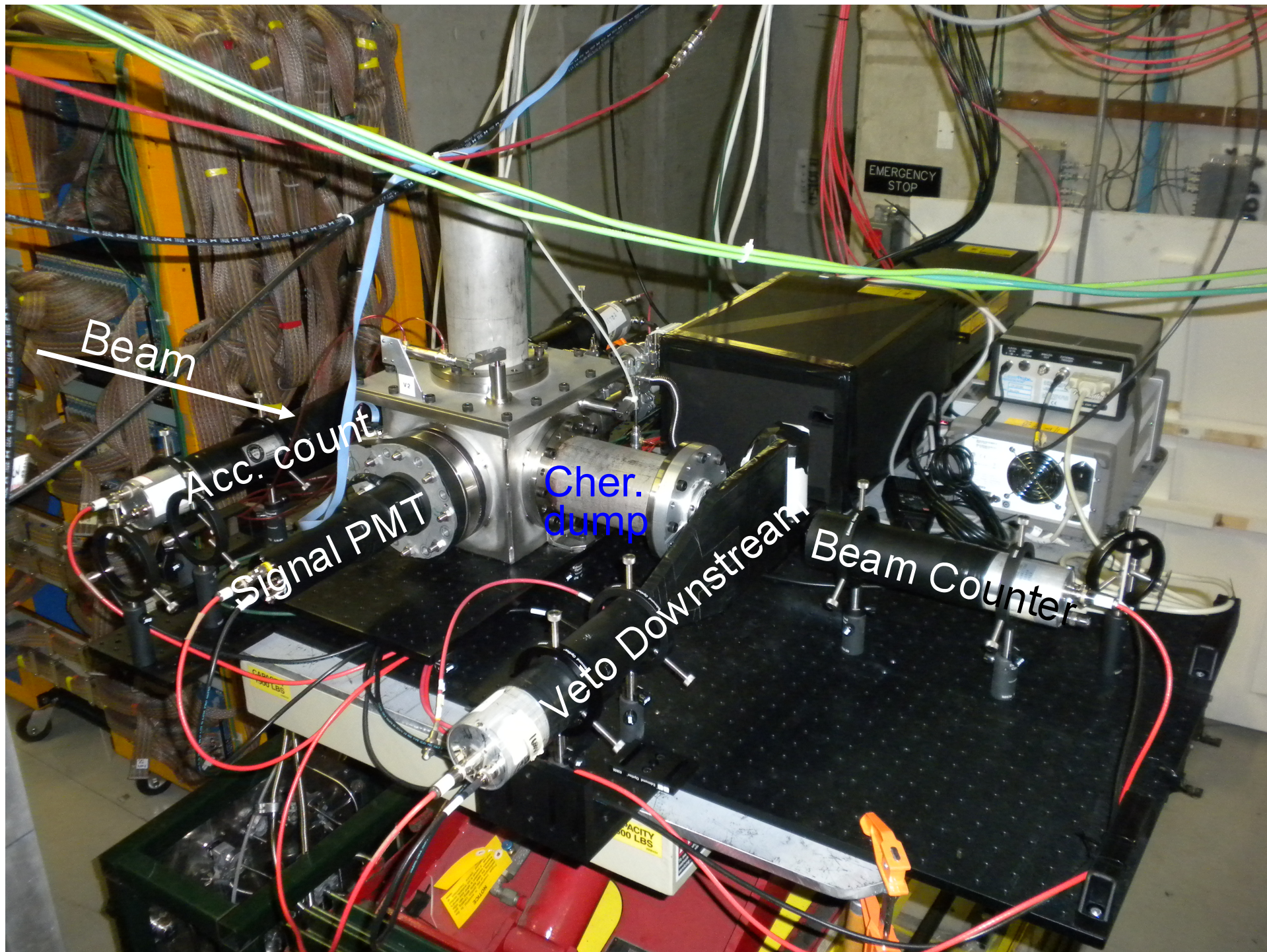


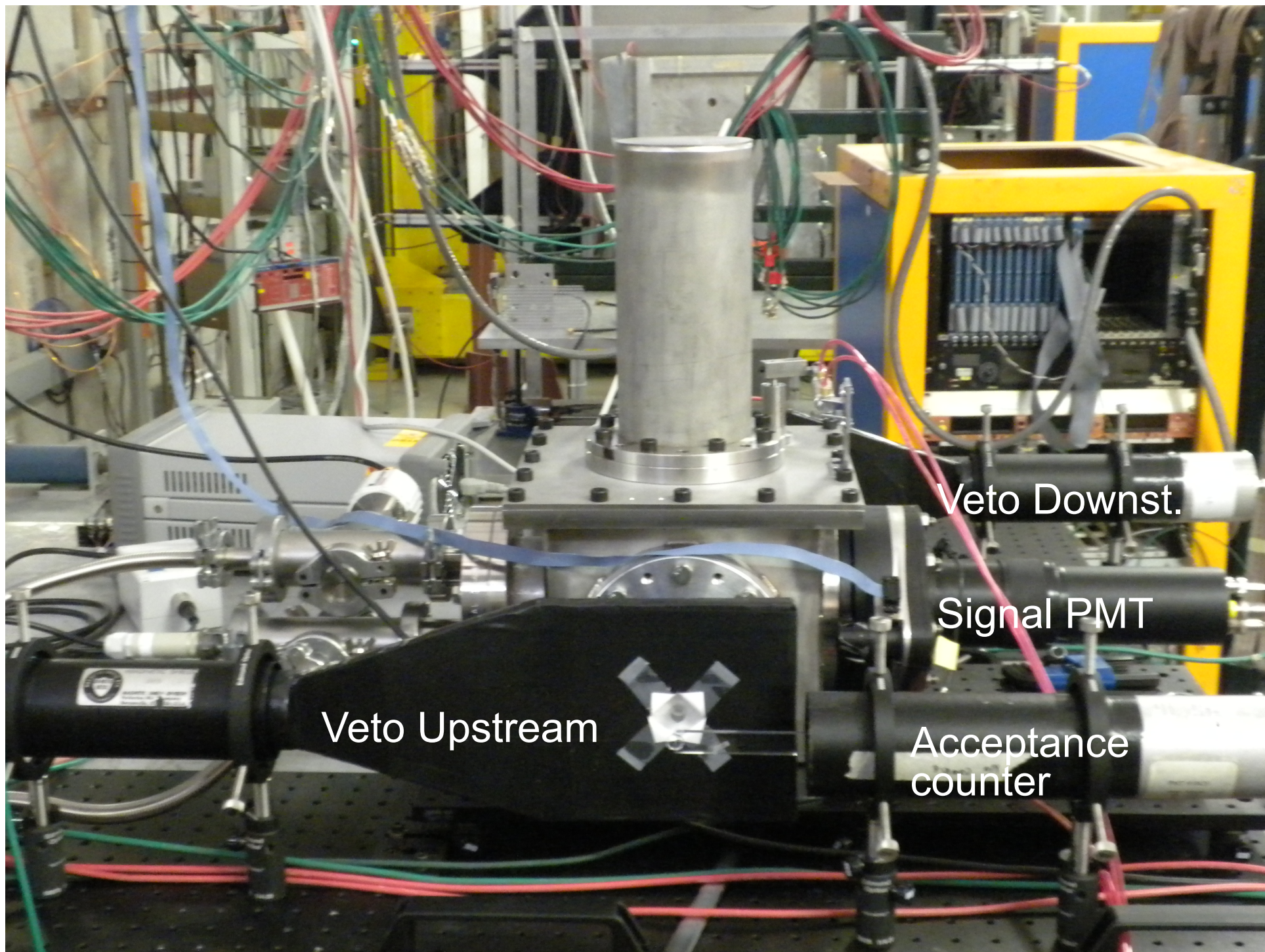
AIRFLY at the Fermilab Test Beam



- High energy up to 120 GeV
- Well defined beam: single particle trigger and geometry
- Wide range of particles type and intensity (p , e , μ , π)

- Absolute calibration with two independent methods: Cherenkov and laser light





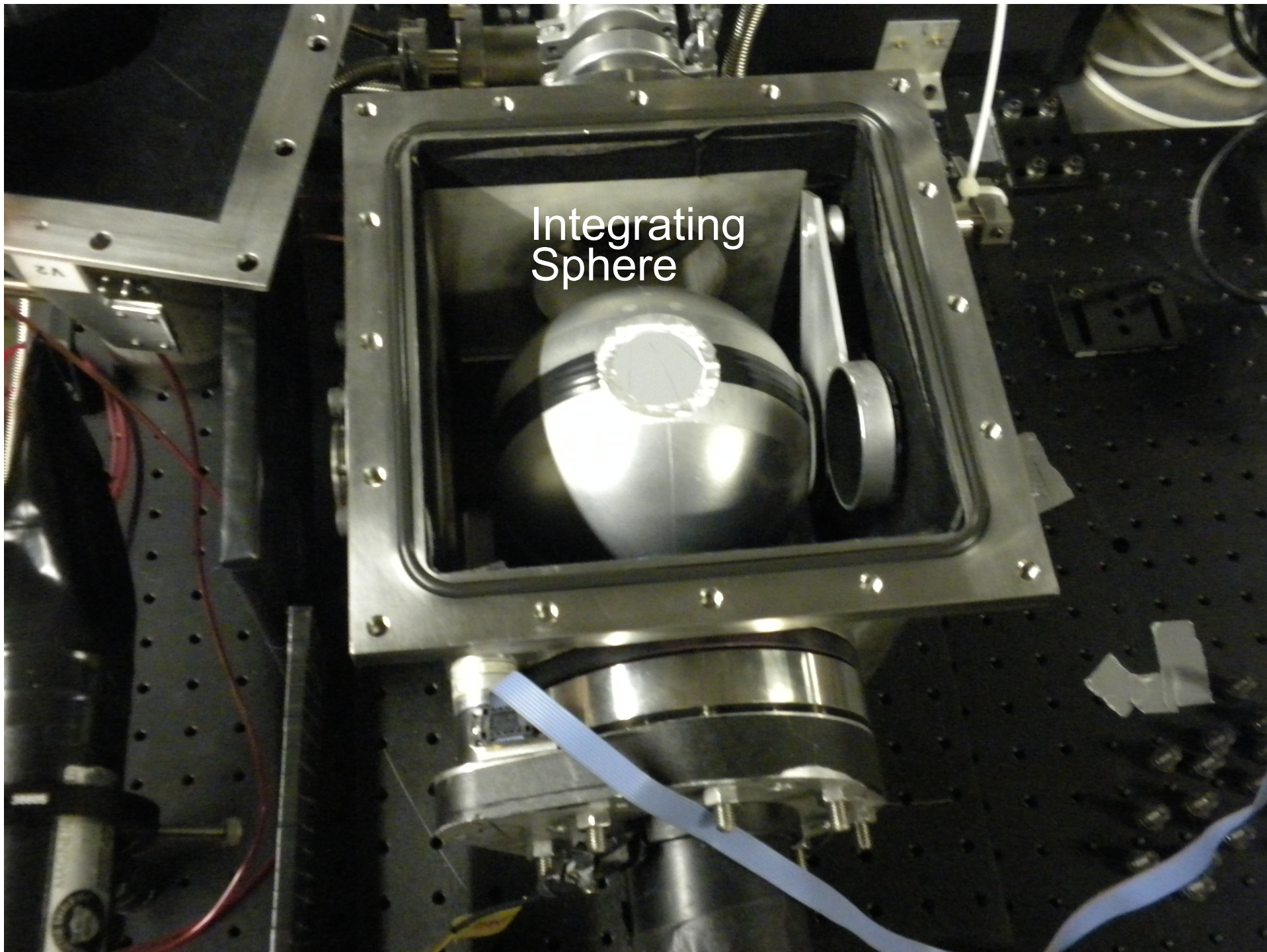
Veto Downst.

Signal PMT

Veto Upstream

Acceptance
counter

Integrating
Sphere



Why use an Integrating Sphere?

- Lambertian light output independent of input light distribution

Fluorescence isotropic –

Cherenkov highly beamed

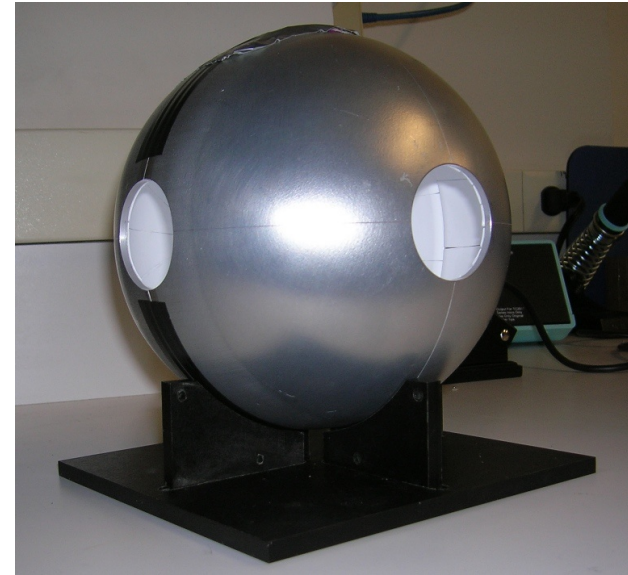
- Depolarize light

- Fluorescence unpolarized –

Cherenkov polarized

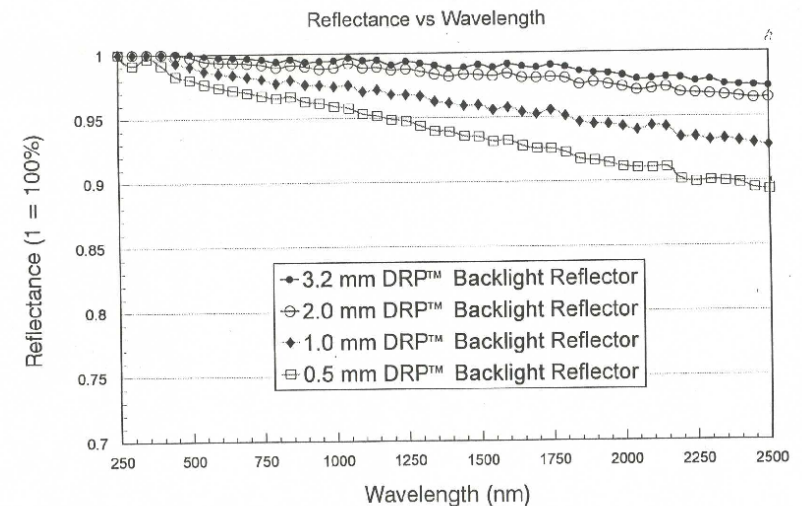
➡ **no syst. from angular or polarization dependent efficiency of PMT+337 nm filter**

- Collect photons over $\sim 4\pi$
Improve signal/bkg ratio



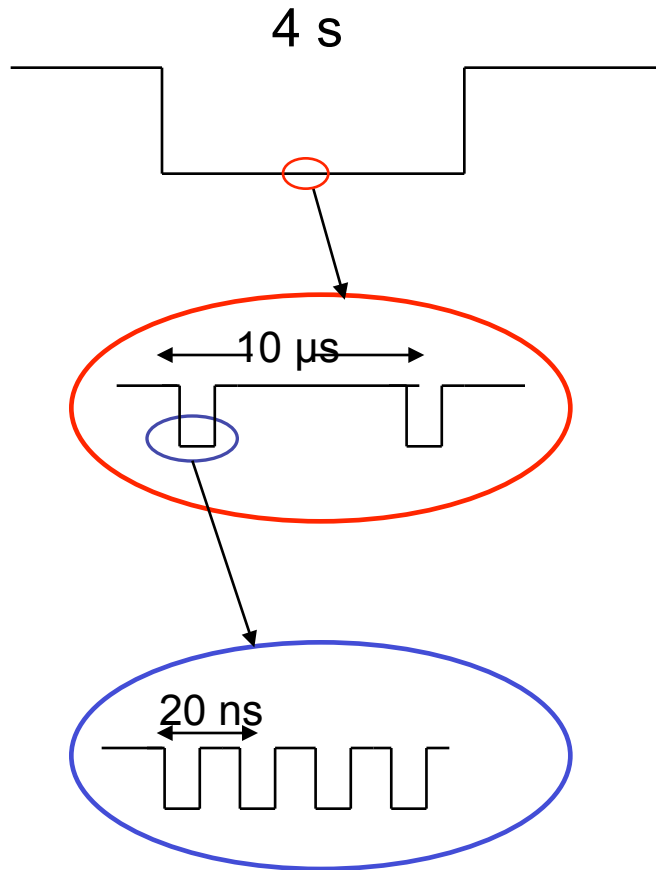
Highly diffusive PTFE

DRP™ BACKLIGHT REFLECTORS AT COMPARATIVE THICKNESS LEVELS



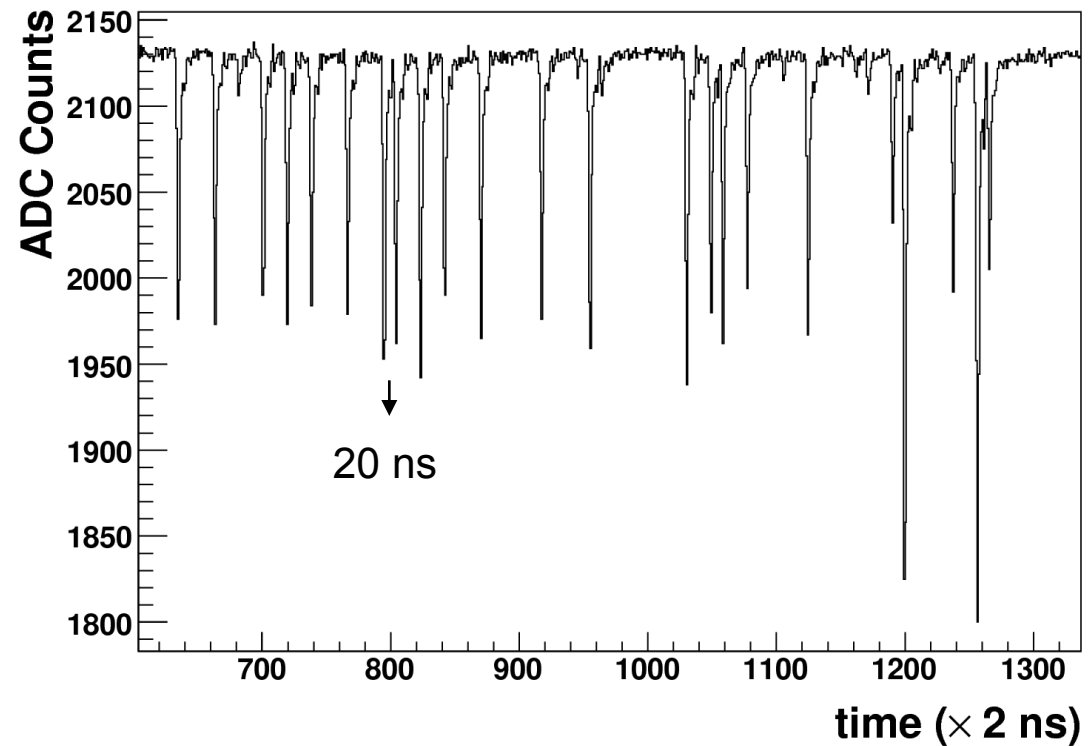
W. L. GORE & ASSOCIATES, II
100 AIRPORT ROAD • P.O. BOX 1010 • ELKTON, MD 21922-1010
PHONE 410/392-3200
FAX 410/392-4817

Single particle triggering



1 spill / min

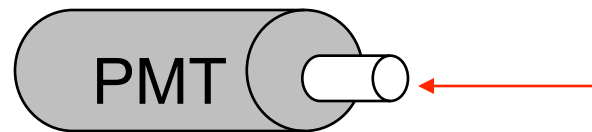
Several 10^5 particles/spill



Up to 70 bunches

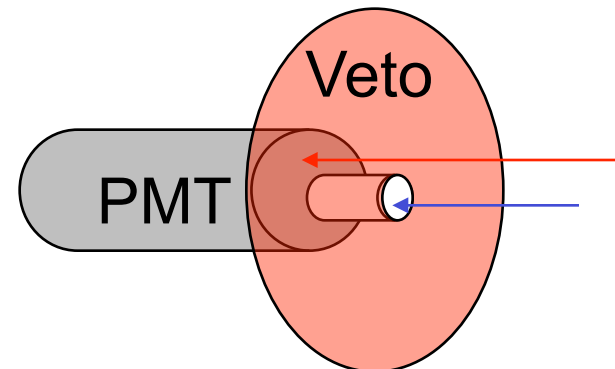
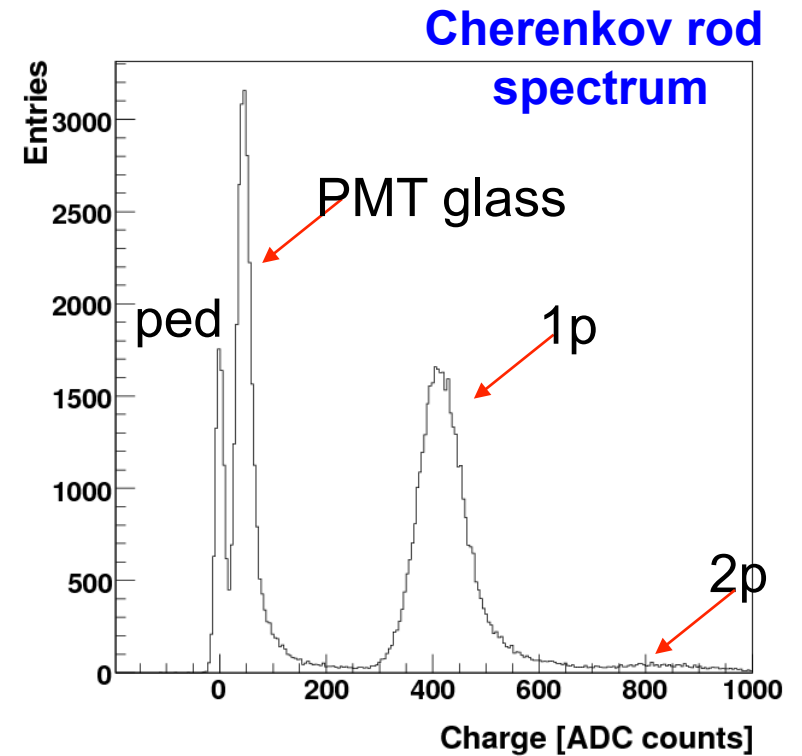
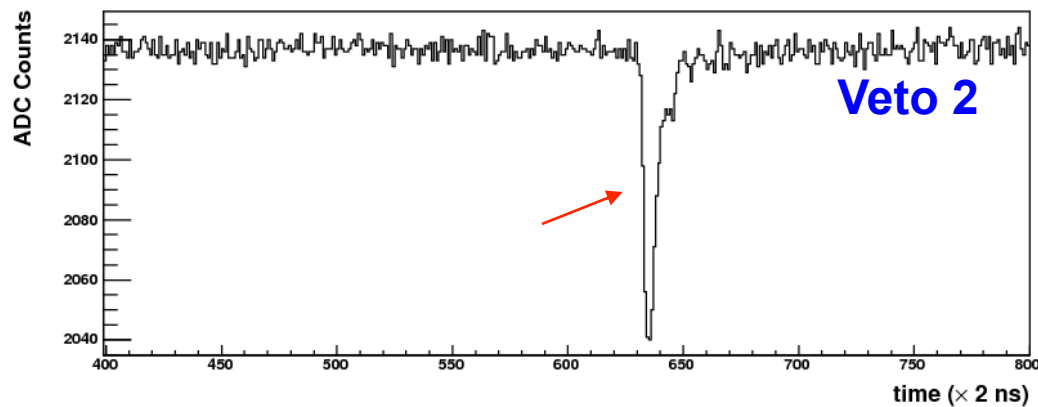
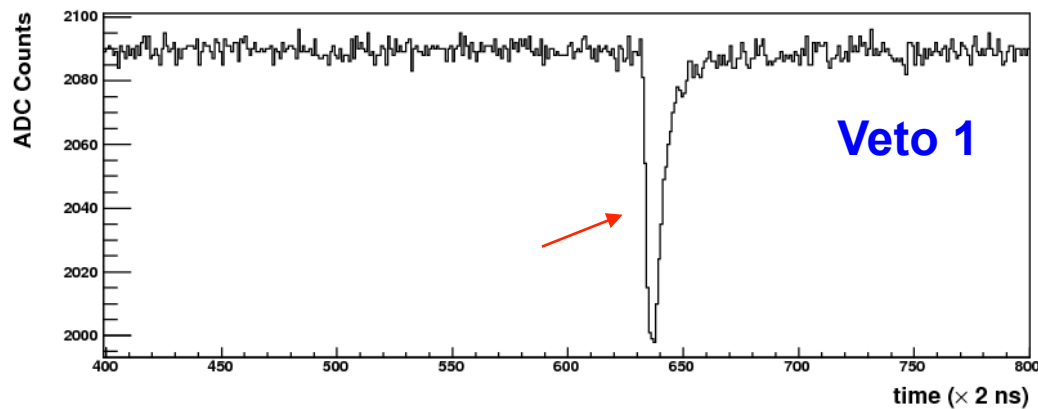
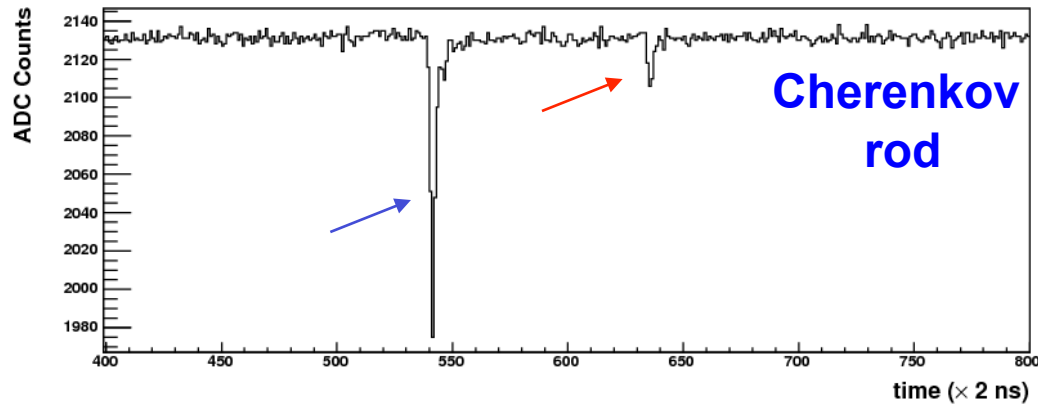
Need single particle resolution, fast detector

Cherenkov, FADC 500 MHz



UV transparent acrylic rod

Single particle triggering

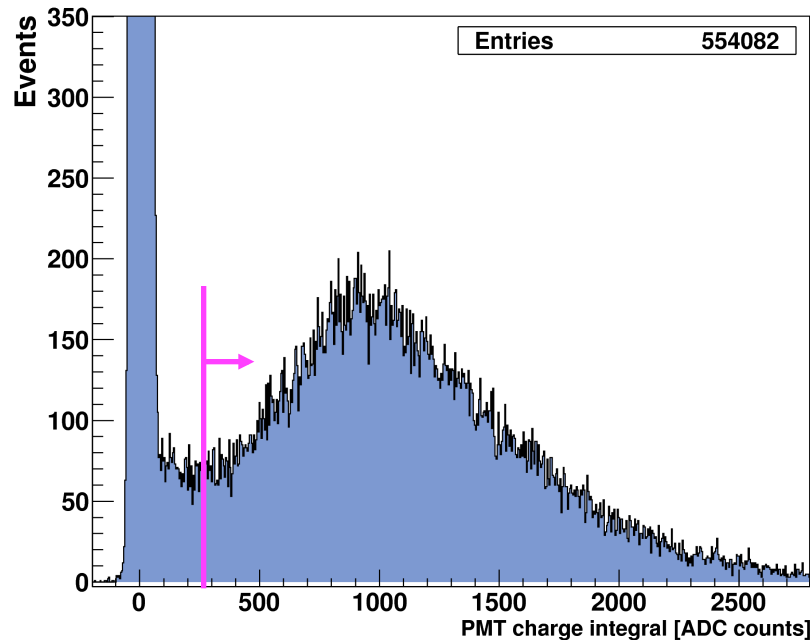


Very clean single particle selection

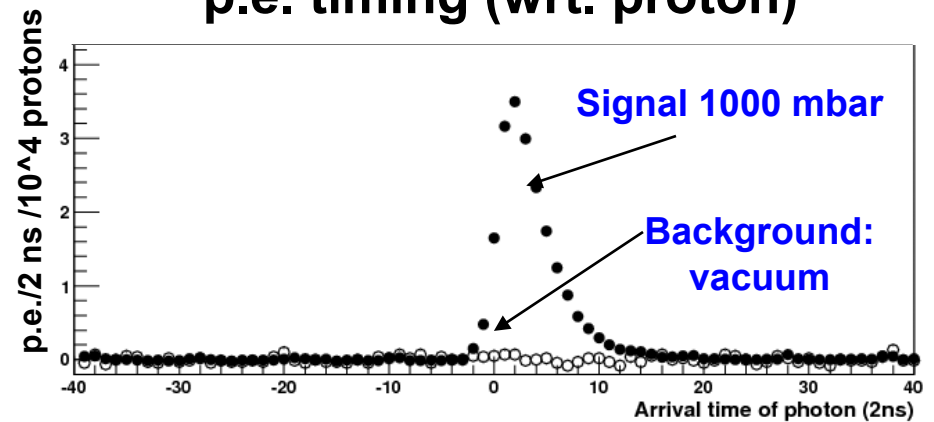
Fluorescence Measurement in N₂

120 GeV protons

Single p.e. spectrum



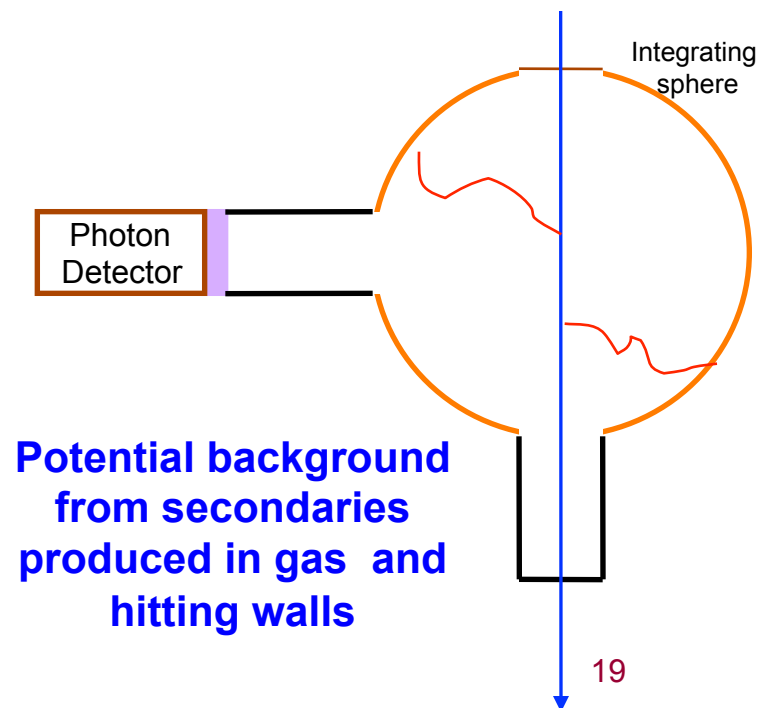
p.e. timing (wrt. proton)



$$S_{FL}(N_2) = (20.05 \pm 0.11) \cdot 10^{-4} \text{ p.e./proton}$$

$$S_{FL}(\text{vac}) = (0.48 \pm 0.09) \cdot 10^{-4} \text{ p.e./proton}$$

Counting experiment: select (single, clean) protons and then count the photons they produced.

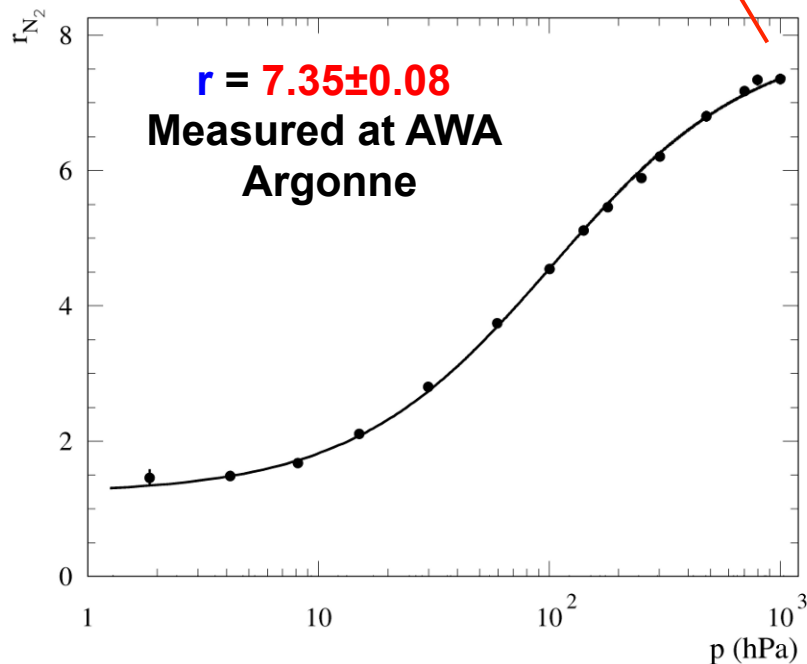


Fluorescence Measurement in N₂ and air

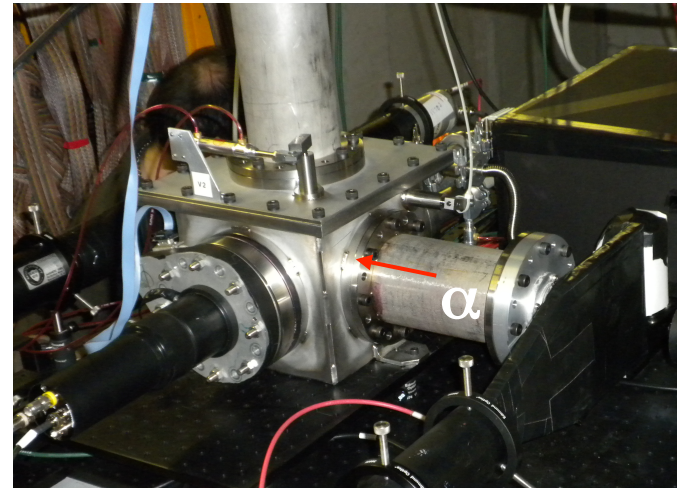
$$S_{FI}(N_2) - S_{FI}(air) = (FI_{N_2} + Bkg) - (FI_{air} + Bkg)$$

Same background from
secondaries produced
in air and N₂

$$= FI_{N_2} - FI_{air} = FI_{N_2} (1 - 1/r)$$



$r = 7.45 \pm 0.08$ measured
in the Fermilab
apparatus with an
alpha source



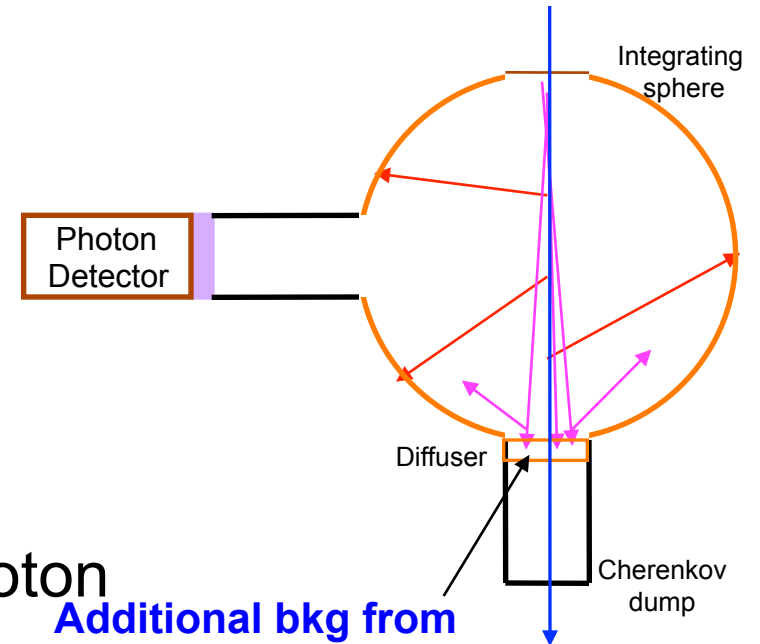
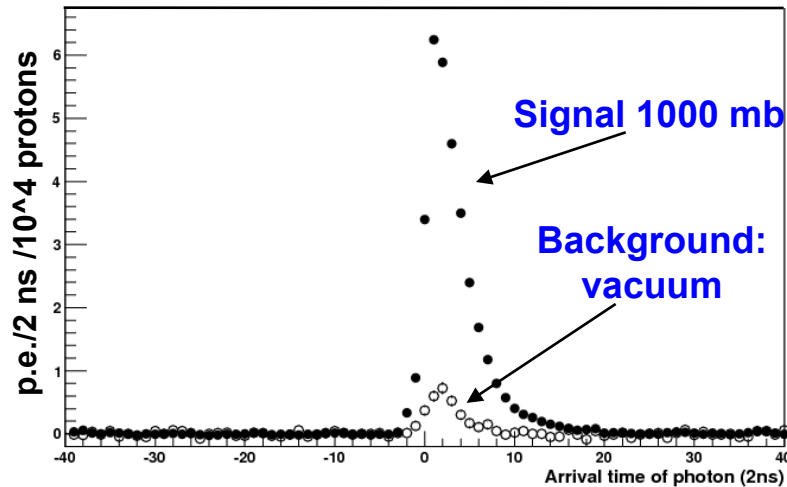
$$FI_{N_2} = (19.44 \pm 0.15) 10^{-4} \text{ p.e./proton}$$

$$Bkg = (0.61 \pm 0.08) 10^{-4} \text{ p.e./proton}$$

1% statistical unc.

consistent with vacuum
bkg only 3% of signal

Cherenkov Measurement in N₂ and air



$$S_{\text{Ch}}(\text{N}_2) = (32.89 \pm 0.15) 10^{-4} \text{ p.e./proton}$$

$$= \text{Ch}_{\text{N}_2} + \text{Fl}_{\text{N}_2} + \text{Bkg} + \text{Bkg}_{\text{Ch}}$$

Additional bkg from
Cherenkov of beam
in DRP (teflon)

$$\text{Bkg}_{\text{Ch}} = (2.57 \pm 0.13) 10^{-4} \text{ p.e./beam particle}$$

Bkg ~10% of Fl+Ch
Measured in vacuum, He

$$\text{Ch}_{\text{N}_2} = (10.27 \pm 0.23) 10^{-4} \text{ p.e./beam particle}$$

2% statistical unc.

$$\text{Ch}_{\text{air}} = (10.28 \pm 0.25) 10^{-4} \text{ p.e./beam particle}$$

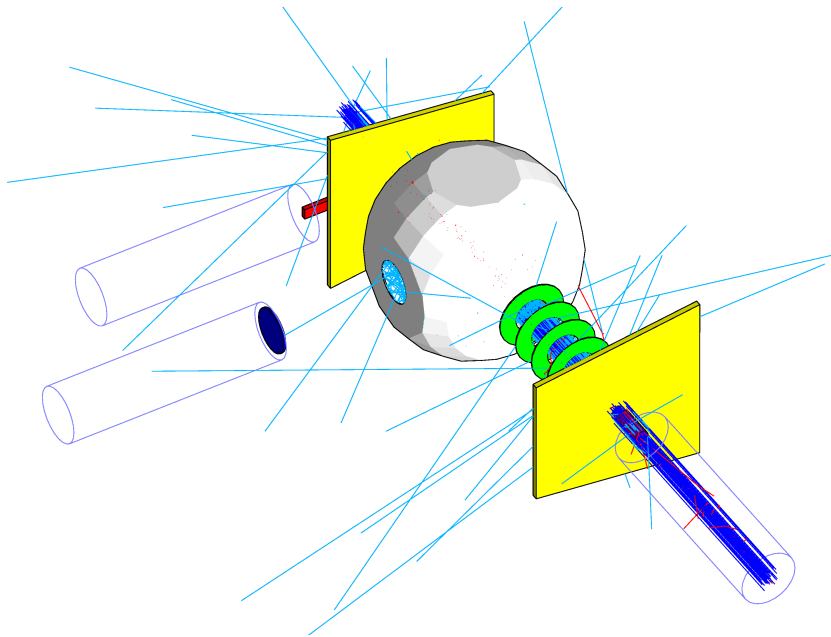
Air ~ N₂ 21

Absolute Yield of 337 nm line

Fluorescence to Cherenkov ratio

$$R_{N2} = \frac{FI_{N2}}{Ch_{N2}} = 1.893 \pm 0.045$$

Preliminary



Geant4 MC simulation of the experiment.

All individual components simulated according to measurement: wavelength and angular dependence of 337 nm filter transmittance, sphere transmission, PMT QE, relative intensities of fluorescence bands, etc.

Given an absolute fluorescence yield in the simulation, $(Y_{air})_{MC}$, we obtain the expected Fluorescence/Cherenkov ratio, $(R_{air})_{MC}$

$$Y_{air} = \frac{(R_{N2} / r)}{(R_{air})_{MC}} \quad (Y_{air})_{MC} = 5.60 \pm 0.13_{stat} \text{ photons}_{337}/\text{MeV}$$

2.4 % statistical unc., systematic discussed later

Currently used in Auger: 5.05 photons/MeV (Nagano)

A few of many checks

- Measurement reproducibility:

2 test beam periods, each 2 weeks. Apparatus taken apart and mounted again. Very stable results. Gas purity stable.

- Is DRP scintillating?

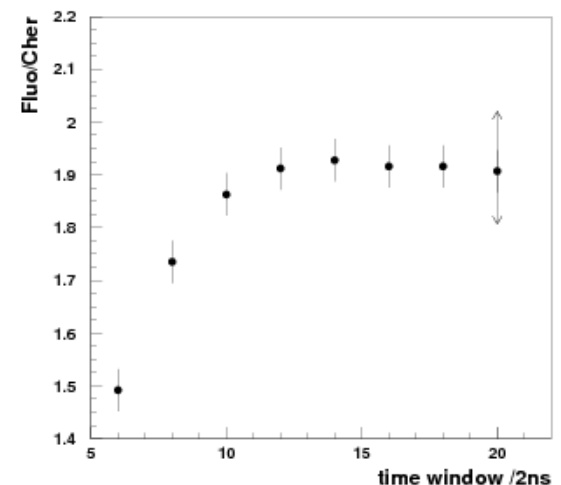
Checked with alphas hitting the sphere under vacuum. No scintillation observed.

- Bkg from secondaries produced in the DRP port in Cherenkov mode?

$S_{\text{Ch}}(\text{N}_2) - S_{\text{Ch}}(\text{air}) = S_{\text{Fl}}(\text{N}_2) - S_{\text{Fl}}(\text{air})$, thus no bkg from secondaries produced in the DRP port

- Fluorescence/Cherenkov stability

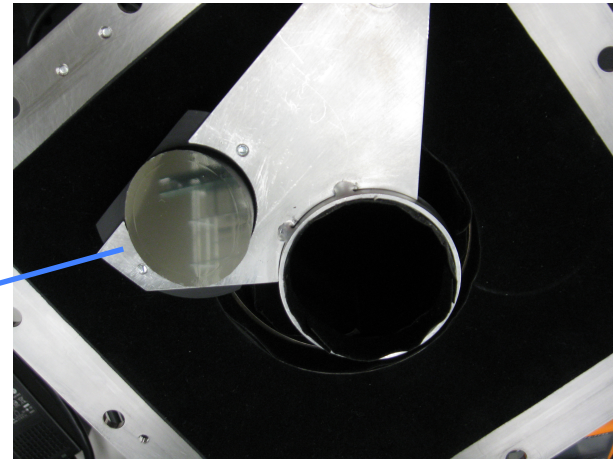
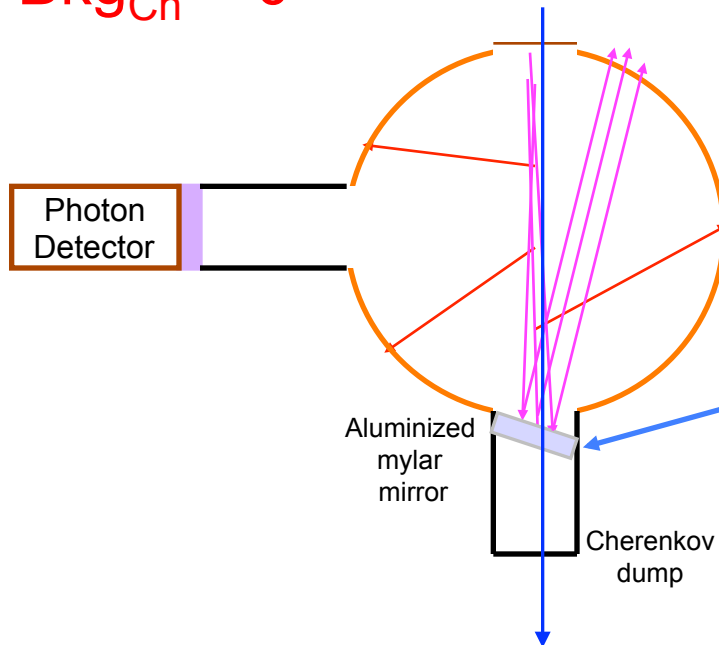
Eg. with respect to time integration of signal



Cross-Check of FI/Ch

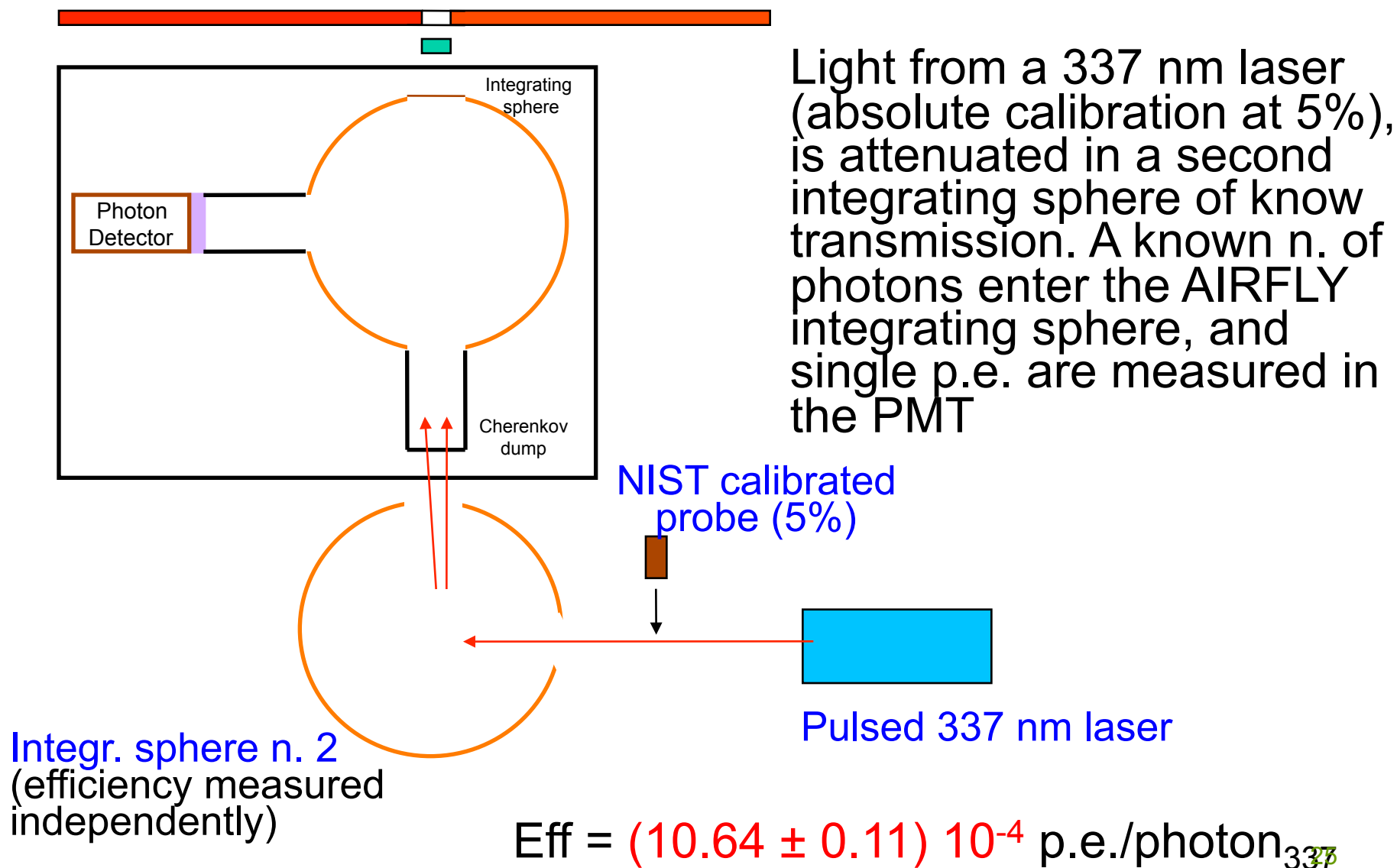
Use an aluminized mylar mirror to bounce Cherenkov light back into the sphere. Eliminate bkg from beam Cherenkov in the DRP

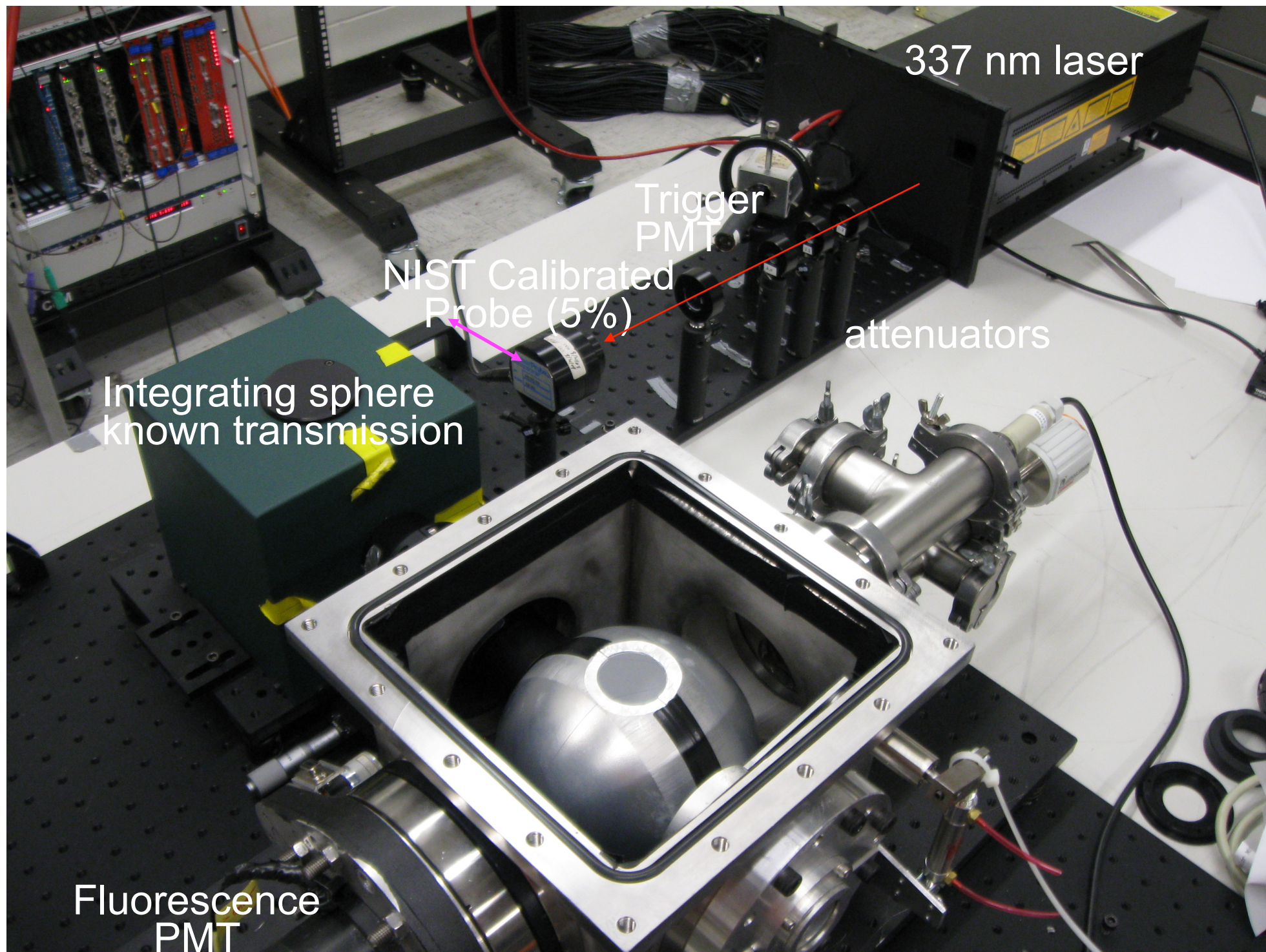
$$\text{Bkg}_{\text{Ch}} = 0$$



Independent Y_{air} measurement, consistent within 2% (4% stat. uncertainty)

Absolute calibration with 337 nm laser





Absolute Yield of 337 nm line

Laser Calibration

$$L_{N2} = \frac{F_{I_{N2}}}{Eff} = (1.757 \pm 0.023) \text{ photons}_{337}/\text{proton}$$

A Geant4 MC simulation of the laser calibration is performed and $(Eff)_{MC}$ is obtained

$$Y_{air} = \frac{(L_{N2} / r)}{(L_{air})_{MC}} \quad (Y_{air})_{MC} = 5.56 \pm 0.07_{stat} \text{ photons}_{337}/\text{MeV}$$

1.3 % statistical unc.

Consistent with the Fluorescence to Cherenkov ratio measurement

NOTE:

In situ laser calibration during test beam. Laser calibration reproduced in the lab at 1% level. Different probes consistent within 3%

Fluorescence data sample independent of the FI/Ch measurement

337 nm laser line vs Cherenkov light integrated over a 10 nm FWHM interference filter. Different systematic

Systematic Uncertainties

Fluorescence to Cherenkov ratio

- Reflectivity of the sphere ~ 1%
- PMT quantum efficiency ~ 1.5%
- Monte Carlo statistics ~ 1%
- N₂/Air ratio ~ 1%
- Geometry ~ 1%
- Filter transmittance ~ 2%
- Background subtraction ~ 1%
- E_{dep} model in MC (% level, large sphere
Ongoing check with F.Arqueros)
-

Preliminary

Total systematic uncertainty expected < 5%

Systematic Uncertainties

Laser Calibration

- Laser probe calibration	~ 5%	<i>Preliminary</i>
- Transmission sphere n. 2	~ 2%	
- Monte Carlo statistics	~ 1%	
- N ₂ /Air ratio	~ 1%	
- Geometry	~ 1%	
- Background subtraction	~ 1%	
- E _{dep} model in MC	(% level, large sphere)	
-		

NOTE:

Largest systematic (5%) from calibrated probe (laser pulse). Currently working to reduce uncertainty to 2% by a DC NIST calibrated probe.

Fluorescence produced by secondary electrons: high energy beam particle not important. We confirmed by measuring fluorescence with 32 GeV pions and 8 GeV electrons.

Summary

- AIRFLY has performed a precise measurement of the absolute fluorescence yield of the 337 nm line. Two independent calibration methods have been used, Cherenkov and laser light, giving compatible results. The preliminary results are consistent within the larger uncertainties of other experiments. We expect a final measurement with a total uncertainty of $\sim 5\%$.
- Together with AIRFLY measurements of the air fluorescence spectrum and its pressure, temperature and humidity dependence, the total uncertainty on the energy scale of UHECR will be reduced to $\sim 5\%$.